

# 1 Solving Poisson Equation with Slowing-down Equilibrium Distribution for 2 Global Gyrokinetic Simulation

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6 Abstract: A generalized multi-point average method has been developed for gyrokinetic  
7 Poisson equation with slowing-down equilibrium distribution and verified for its accuracy in  
8 the long and short wavelength limits, which forms an important basis for global gyrokinetic  
9 simulation of low frequency drift Alfvénic turbulence in burning plasmas.

10 Keyword: Gyrokinetics, Poisson Solver, Slowing-down Distribution

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## 12 I. Introduction

13 In gyrokinetic particle simulation, the difference between particle distribution and gyro-  
14 center distribution leads to the double gyro-average of potential field, which is manifested in  
15 polarization density [1,2]. The polarization density, which is essential for gyrokinetic Poisson  
16 equation, depends critically on the phase structure of the equilibrium particle distribution,  
17 especially for short wavelength modes. Global gyrokinetic simulations are crucial for studying  
18 many important physics issues in magnetic fusion plasmas, such as turbulent transport scaling  
19 and turbulence spreading [3, 4, 5, 6, 7]. With the advent of burning plasmas, the alpha  
20 particles would inevitably excite Alfvénic turbulence with a slowing-down distribution via  
21 electron-alpha collisions [8, 9, 10, 11, 12, 13]. A more accurate global gyrokinetic Poisson  
22 solver with slowing-down background distribution is desirable for simulating alpha particle  
23 physics in burning plasmas and NBI heating scenarios.

24 Some recent researches have used so-called equivalent Maxwellian distribution, whose  
25 temperature or second order velocity moment is the same as slowing-down distribution [14]  
26 to simulate the alpha particle physics, which may be valid for a number of physics scenarios  
27 that depends weakly on the phase space structures of the equilibrium distribution. And

28 Some other work [15], though correctly consider velocity space derivatives in the gyrokinetic  
 29 equations, uses flux-tube method to avoid tackling global spatial dependence of the  
 30 gyrokinetic Poisson equation. In this paper, a novel global method based on multi-point  
 31 average [2,16] to solve this equation is developed to adapt the slowing-down equilibrium  
 32 distribution, and the accuracy of this new method is verified in the long and short wavelength  
 33 limits. This method is essential for global gyrokinetic simulation to investigate alpha particle  
 34 physics in the burning plasmas with the advent of ITER operation.

35 The rest of this article is structured as the following: Sec. II describes how to derive  
 36 polarization density in gyrokinetics with slowing-down particle distribution; Sec. III shows our  
 37 numerical scheme to solve gyrokinetic Poisson equation with the slowing-down equilibrium  
 38 distribution; and numeric verification is provided in Sec. IV for the accuracy of this new scheme;  
 39 The numerical results are summarized in Sec. V.

## 40 II. Gyrokinetic Poisson Equation

41 The gyrokinetic-Maxwellian system can be expressed as

$$42 \frac{\partial}{\partial t} \bar{f} + (v_{\parallel} \mathbf{b} + \mathbf{v}_d + \mathbf{v}_E) \cdot \nabla \bar{f} - \frac{\mathbf{B}^*}{m_s B_0} \cdot (\mu_s \nabla B + q(\nabla \langle \delta \phi_{gc} \rangle + \mathbf{b} \partial_t \langle \delta A_{gc\parallel} \rangle)) \frac{\partial}{\partial v_{\parallel}} \bar{f} = 0, \quad (1)$$

$$43 q_i \delta n_i = e \delta n_e, \quad (2)$$

$$44 \nabla_{\perp}^2 \langle \delta A_{gc\parallel} \rangle = \mu_0 \sum_{s=i,e} q_s \delta u_{\parallel s} \quad (3)$$

45 where the gyro-center coordinates  $(\mathbf{X}, v_{\parallel}, \mu, \zeta)$  are used,  $\mathbf{v}_d = \mathbf{b} \times (\mu \nabla B + m_s v_{\parallel}^2 \boldsymbol{\kappa}) / (q_s B)$   
 46 is magnetic drift for the guiding centers,  $\mathbf{v}_E = -\nabla \langle \delta \phi_{gc} \rangle / B$  is the gyro-averaged  $\mathbf{E} \times \mathbf{B}$   
 47 drift, and  $\langle \dots \rangle$  represents gyrophase average.  $\delta \phi_{gc}$  is the perturbed potential at the gyro-  
 48 center, which is defined as  $\delta \phi_{gc}(\mathbf{X}; \mu, \zeta, t) \equiv \exp(\boldsymbol{\rho} \cdot \nabla) \delta \phi = \delta \phi(\mathbf{x}, t)$  and the gyro-radius  $\boldsymbol{\rho}$   
 49 is defined as  $\boldsymbol{\rho} = \mathbf{b} \times \mathbf{v}_{\perp} / \Omega_s$ , where the gyrofrequency  $\Omega_s = q_s B / m$  and magnetic moment  
 50  $\mu_s = m_s v_{\perp}^2 / 2B$ . And  $\delta A_{gc\parallel}$  is perturbed parallel vector potential at the gyro-center.  $\mathbf{B}^* =$   
 51  $\mathbf{B}_0 + m_{\alpha} v_{\parallel} \nabla \times \mathbf{b} / q_{\alpha} + \delta \mathbf{B}$ . Suppose that the distribution function  $f$  can be decomposed into  
 52 an equilibrium component  $f_0$  and a perturbed component  $\delta f$ , i.e.,  $f = f_0 + \delta f$ , the  
 53 perturbed density  $\delta n_s$  can be calculated by  $\delta n_s = \int d^3 v \delta f_s$  and perturbed parallel fluid  
 54 velocity  $\delta u_{\parallel s}$  can be calculated by  $\delta u_{\parallel s} = \int d^3 v v_{\parallel} \delta f_s$ . As is shown in Eq. (2), the quasi-

55 neutrality for the fluctuating densities  $\delta n_s$  is used to solve for the electrostatic potential  $\delta\phi$ ,  
 56 which is valid for wavelengths longer than the Debye length. In many cases, the adiabatic  
 57 response is assumed for electrons due to their fast parallel motion, i.e.,  $\delta n_e = en_0\delta\phi/T_e$ . In  
 58 the last, the gyrokinetic parallel Ampère's law [17] in Eq. (3) is used for solving the vector  
 59 potential  $\delta A_{\parallel}$  and the gyrokinetic-Maxwellian system is closed.

60 Employing the Lie transform method [18], we can calculate the total ion density  $n_i$  by  
 61 integrating over the velocity space a peculiar distribution function, which is generated by  
 62 pulling back the gyro-center distribution function  $\bar{f}$  into the particle coordinate space:

$$63 \quad n_0 + \delta n_i = \int d^3v \exp(-\boldsymbol{\rho} \cdot \nabla) \left( 1 + \frac{q_i}{B} \delta\tilde{\phi}_{gc} \frac{\partial}{\partial\mu} \right) \bar{f}$$

64 Here  $\delta\tilde{\phi}_{gc} \equiv \delta\phi_{gc} - \langle \delta\phi_{gc} \rangle$  is the gyrophase dependent part of  $\delta\phi_{gc}$ . It is found that the  
 65 perturbed ion density  $\delta n_i$  can be separated into a perturbed guiding center density and a  
 66 polarization density, i.e.,  $\delta n_i = \delta n_{i,gc} + n_{i,pol}$ , where the guiding center density  $\delta n_{i,gc}$  is  
 67 defined as  $\delta n_{i,gc} = \int d^3v \exp(-\boldsymbol{\rho} \cdot \nabla) \delta\bar{f}$ , and the polarization density  $n_{i,pol}$  is associated  
 68 with the fluctuating electric field  $\delta\phi$  through

$$69 \quad n_{i,pol} = \int d^3v \exp(-\boldsymbol{\rho} \cdot \nabla) \left( \frac{q_i}{B} \delta\tilde{\phi}_{gc} \frac{\partial}{\partial\mu} f_0 \right). \quad (4)$$

70 As for the parallel Ampère's law of Eq. (3), one can find that there's no explicit dependence  
 71 on the perturbed electromagnetic field in the right-hand side of the equation, thus no extra  
 72 modifications are needed when the equilibrium distribution function is changed from  
 73 Maxwellian to slowing-down.

74 To solve the gyrokinetic Poisson equation, a proper numerical algorithm is needed to deal  
 75 with the guiding center transformation  $\exp(-\boldsymbol{\rho} \cdot \nabla)$  and gyrophase average in Eq. (4). The  
 76 four-point gyro-average method has been invented [16] to solve the gyrokinetic Poisson  
 77 equation in real space when the equilibrium distribution  $f_0$  is Maxwellian, which enables a  
 78 global gyrokinetic particle simulation [16,18]. Here we modify the original four-point average  
 79 method to accommodate a slowing-down equilibrium distribution in order to investigate the  
 80 self-consistent turbulent transport physics involving alpha particles, i.e.,  $f_0 = f_{std}$ .

81 The slowing-down distribution  $f_{std}$  is the steady state solution to the collisional scattering  
 82 for an isotopic particle source with a large birth speed  $v_0$ , i.e., the alpha particle speed  
 83 produced by the thermal nuclear fusion, e.g.,  $\frac{1}{2} m_{\alpha} v_0^2 = 3.5 \text{ Mev}$  for a typical D-T fusion. It is

84 discovered that  $f_{slid}$  can be explicitly expressed as [19]:

$$85 \quad f_{slid}(v) = \frac{n}{4\pi I_1} \frac{H(v_0 - v)}{v^3 + v_c^3}, \quad (5)$$

86 where  $v_c \equiv \left( \frac{3\sqrt{\pi} m_e}{4n_e} \sum_i \frac{q_i^2 n_i}{e^2 m_i} \right)^{1/3} v_{th,e}$  is the critical speed at which the electron drag is  
 87 comparable to the thermal ion drag, and  $I_1 = \frac{1}{3} \ln(1 + v_0^3/v_c^3)$  is an auxiliary function for the  
 88 purpose of normalization. A “temperature” can be defined for the slowing-down distribution  
 89 as the 2<sup>nd</sup> velocity moment of the distribution function, which is similar to that of the  
 90 Maxwellian:

$$91 \quad \frac{3}{2} n T_{slid} = \int d^3v \frac{1}{2} m_i v^2 f_{slid}(v) \equiv \frac{1}{2} n m_i v_c^2 \frac{I_2}{I_1} \quad (6)$$

92 with  $I_2 = \frac{v_0^2}{2v_c^2} - \frac{1}{6} \left( \frac{\pi}{\sqrt{3}} - 2\sqrt{3} \arctan \frac{1-2v_0/v_c}{\sqrt{3}} - \ln \frac{(1+v_0/v_c)^3}{1-v_0^3/v_c^3} \right)$ . We note that for the fusion born  
 93 alpha particles in a 10keV 50%-50% D-T plasma,  $v_c/v_0 \approx 0.3$ ,  $T_{slid} \approx 1.28 m_i v_c^2$ .

94 Consider the Fourier representation of the perturbed potential  $\delta\phi = \sum_k \delta\phi_k \exp(i\mathbf{k} \cdot \mathbf{x})$  in  
 95 Eq. (4), and choose  $f_0$  as the slowing-down distribution, then we can obtain:

$$96 \quad n_{i,pol} = - \sum_k \int d^3v (1 - \exp(-i\mathbf{k} \cdot \boldsymbol{\rho})) J_0(k_\perp \rho) \delta\phi_k \lambda f_0 \exp(i\mathbf{k} \cdot \mathbf{x})$$

$$97 \quad = - \sum_k (c_0 - \Gamma_0(k_\perp \rho_c)) \frac{n_i}{T_i} q_i \delta\phi_k \exp(i\mathbf{k} \cdot \mathbf{x}), \quad (7)$$

98 where  $\rho_c = v_c/\Omega_i$ ,  $\lambda = \frac{T_i}{m_i n_i} \left( \frac{3v}{v^3 + v_c^3} + \frac{\delta(v_0 - v)}{v} \right)$ ,  $c_0 = \int d^3v \lambda f_0 \equiv \frac{T_i}{m_i v_c^2} \frac{I_3}{I_1}$  and  $I_3 = \frac{1}{6} \left( \frac{\pi}{\sqrt{3}} - \right.$   
 99  $\left. 2\sqrt{3} \arctan \frac{1-2v_0/v_c}{\sqrt{3}} + \ln \frac{(1+v_0/v_c)^3}{1-v_0^3/v_c^3} \right)$ .  $J_0 = J_0(k_\perp \rho_i) = \langle \exp(i\mathbf{k} \cdot \boldsymbol{\rho}) \rangle$  is zeroth order Bessel  
 100 function,, and  $\Gamma_0(k_\perp \rho_c)$  is defined as

$$101 \quad \Gamma_0(k_\perp \rho_c) \equiv \int d^3v \exp(-i\mathbf{k} \cdot \boldsymbol{\rho}) J_0(k_\perp \rho) \lambda f_0 = \int dv_\perp v_\perp J_0^2(k_\perp v_\perp / \Omega_i) \tilde{f}(v_\perp), \quad (8)$$

102 which can be considered as the expectation of  $\lambda$  weighted by the equilibrium distribution  $f_0$   
 103 after double gyroaveraging due to the back and forth transformation between particle  
 104 position and gyrocenter position, where  $\tilde{f}(v_\perp) = \int d\mathbf{v}_\parallel \lambda f_0$ . In the case for  $f_0$  to be  
 105 Maxwellian, the function of  $\Gamma_0$  can be calculated analytically [1], i.e.,  $\Gamma_0 = I_0(b) e^{-b}$ , with  $b =$   
 106  $k_\perp^2 \rho_c^2$  and  $\rho_c = \sqrt{T/m}$ . Unlike the Maxwellian equilibrium case,  $\Gamma_0$  do not have a simple  
 107 analytic expression in Fourier space when  $f_0$  is slowing-down and has to be evaluated  
 108 numerically. In principle, one can solve Eq. (2) using Eq. (7) and (8) in Fourier space. With these  
 109 newly defined functions, the gyrokinetic Poisson equation can be written in a dimensionless

110 form

$$111 \quad \left(\frac{q_i}{T_i} c_0 + \frac{e}{T_e}\right) \delta\phi - \frac{q_i}{T_i} \widetilde{\delta\phi} = \frac{\delta n_{i,gc}}{n_i}, \quad (9)$$

112 where  $\widetilde{\delta\phi}$  has a complicated form in the real space but a neat form in the Fourier space:

$$113 \quad \widetilde{\delta\phi} \equiv \sum_k \Gamma_0(k_\perp \rho_c) \delta\phi_k \exp(i\mathbf{k} \cdot \mathbf{x}). \quad (10)$$

114 However, this Fourier representation is not always valid since it mixes up the configuration  
115 space and velocity space dependences through  $J_0$  term. In reality, the background magnetic  
116 field and perpendicular temperature can vary in real space, and then  $\Gamma_0$  will gain global  
117 spatial dependences. Besides, the Fourier transform approach is more difficult to deal with  
118 realistic tokamak geometry, where no periodicity exists in the radial direction and on many  
119 occasions the global effects have to be considered seriously. For the Maxwellian background  
120 distribution, the four-point gyro-average method has been developed to solve this  
121 gyrokinetic Poisson equation in the real space [1,2]. Here we improve this method by  
122 including the slowing-down background distribution  $f_{sl,d}$  as the equilibrium distribution  $f_0$   
123 in the gyrokinetic Poisson equation, i.e., Eq. (9)

### 124 III. Gyrokinetic Poisson Solver with Slowing Down 125 Distribution

126 The crucial part of implementing this gyrokinetic Poisson solver in the gyrokinetic  
127 simulation is to represent  $\widetilde{\delta\phi}$  in Eq. (10) by the values of  $\delta\phi$  at various field points in the  
128 real space. By numerical interpolation, we note that  $\widetilde{\delta\phi}$  can be expressed as a linear  
129 combination of the  $\delta\phi$  values on a number of nearby grid points and consequently Eq. (9)  
130 is transformed into a discrete matrix form such as  $\mathbf{A} \cdot \mathbf{x} = \mathbf{b}$ , which can then be solved by  
131 many known matrix inversion algorithms.

132 Starting from the definition of  $\widetilde{\delta\phi}$  in its integral form instead of the Fourier form, one  
133 finds that

$$134 \quad \widetilde{\delta\phi} = \int_0^\infty dv_\perp v_\perp \langle \exp(-\boldsymbol{\rho} \cdot \nabla) \langle \delta\phi_{gc} \rangle \rangle \tilde{f}(v_\perp). \quad (11)$$

135 To calculate  $\widetilde{\delta\phi}$  at a grid point  $\mathbf{x}_g$  for a specific  $v_\perp$ , one needs to evaluate the gyroaveraged  
136 function  $\langle \exp(-\boldsymbol{\rho} \cdot \nabla) \langle \delta\phi_{gc} \rangle \rangle$ , which is the average value of  $\langle \delta\phi_{gc} \rangle$  on a ring with radius  $\rho$

137 around  $\mathbf{x}_g$ , as is shown by the dotted circle in Fig. 1. The gyroaveraged quantity  $\langle \delta\phi_{gc} \rangle$  can  
 138 also be calculated by this ring average method, e.g., the value of  $\langle \delta\phi_{gc} \rangle$  at the black triangle  
 139 in Fig. 1 can be calculated by the average value of on solid circle. It is not necessary to actually  
 140 integrate numerically along the whole ring to compute the gyrophase average, which would  
 141 make the gyroaverage process rather time-consuming and expensive. According to [1, 2], A  
 142 selection of four points uniformly distributed on the ring (four-point average method) is  
 143 sufficient to compute the gyroaverage for wavelengths up to  $k_{\perp}\rho_c \sim 2$ . Thus, nine  
 144 neighboring points are required to compute  $\widetilde{\delta\phi}$  on the grid point, as is shown by eight red  
 145 points and the center blue diamond in Fig. 1. In more general geometry, these points required  
 146 for the gyroaverage computation may not lay exactly on the grids, but their values can be  
 147 acquired by a linear interpolation of the nearby grid points. Finally, summing up a few rings  
 148 with different values  $v_{\perp}$  with the weight function  $\tilde{f}(v_{\perp})$  and the relationship between  $\widetilde{\delta\phi}$   
 149 and  $\delta\phi$  on each grid point is found.

150 The remaining issue for evaluating Eq. (8) is how to discretize the  $v_{\perp}$  integral with the weight  
 151 function  $\tilde{f}(v_{\perp})$ . Here we approximate the integral by a weighted summation by choosing a  
 152 few sampling grid points along the  $v_{\perp}$  coordinate. From the definition of  $\widetilde{\delta\phi}$ , one can tell  
 153 that it is equivalent to approximate Eq. (8) by:

$$\begin{aligned}
 156 \quad \Gamma_0(k_{\perp}\rho_c) &= \int dv_{\perp} J_0^2(k_{\perp}v_{\perp}/\Omega_i) \tilde{f}(v_{\perp}) \\
 157 \quad &\approx \sum_j c_j J_0^2\left(k_{\perp}\rho_c \frac{v_{\perp j}}{v_c}\right), \quad (12).
 \end{aligned}$$

154 where  $c_j$  are the summing weights due to  $\tilde{f}(v_{\perp})$  and  $v_{\perp j}$  are the sampling grid points. The  
 155 value pairs of  $(c_j, v_{\perp j})$  are chosen by minimizing the following error function:

$$158 \quad \epsilon = \int_0^a \left( \Gamma_0(x) - \sum_j c_j J_0^2\left(x \frac{v_{\perp j}}{v_c}\right) \right)^2 dx \quad (13)$$

159 Here  $a$  is the maximum value of  $k_{\perp}\rho_c$  that we are interested in. Since low frequency micro-  
 160 turbulence usually peaks around  $k_{\perp}\rho_c < 1$ , it is required that this approximation has a better  
 161 accuracy for long wavelengths or  $k_{\perp}\rho_c \ll 1$ . Considering the Taylor expansion for  $J_0(x)$  and  
 162  $\Gamma_0(x)$  around  $x \sim 0$ , one finds that  $J_0(x) = 1 - x^2/4 + O(x^4)$  and  $\Gamma_0(x) = c_0 - \frac{T}{mv_c^2} x^2 +$   
 163  $O(x^4)$ . Let the first two terms equal to each other:

164 
$$\sum_j c_j = c_0 = \frac{I_1 I_2}{3I_1^2} \quad (14)$$

165 
$$\sum_j c_j \frac{v_{\perp j}^2}{v_c^2} = \frac{2T}{mv_c^2} = \frac{2I_2}{3I_1} \quad (15)$$

166 These two constrains are then used to reduce degree of freedom. In order to minimize  $\epsilon$   
 167 with respect to  $(c_j, v_{\perp j})$ , we use the Nelder-Mead method [19], which is a gradient-free  
 168 iterative optimization algorithm.  $I_{1,2,3}$  are functions of  $v_c/v_0$ , which is chosen to be 0.3 here  
 169 to show numeric result of  $(c_j, v_{\perp j})$ . In the one-velocity-node case, we find that  $c = 1.226$   
 170 with the velocity node  $v_{\perp}/v_c = 1.443$  and the relative error is 3.6% for  $k_{\perp}\rho_c < 0.5$ . When  
 171 using two velocity nodes, we find that  $c_1 = 0.9347$  and  $c_2 = 0.2910$  with the velocity nodes  
 172  $v_{\perp 1}/v_c = 0.8778$  and  $v_{\perp 2}/v_c = 2.510$ , and the relative error is about 3.6% for  $k_{\perp}\rho_c < 1.5$ .  
 173 In the three-velocity-node case, we find that  $(c_1, c_2, c_3) = (0.1186, 0.3881, 0.7190)$  with  
 174  $(v_{\perp 1}/v_c, v_{\perp 2}/v_c, v_{\perp 3}/v_c) = (0.7016, 1.716, 2.984)$ , and the relative error is only 0.46% for  
 175  $k_{\perp}\rho_c < 2$ . The three-velocity-node approximation is compared with the exact value from  
 176 direct numerical integration, as is shown in Fig. 2. Satisfactory accuracy is achieved with a  
 177 relative error less than 0.46% for  $k_{\perp}\rho_c < 2$ , which is sufficient to include most interesting  
 178 finite Larmor radius effects due to the slowing-down alpha particles. We also test for the  
 179 widely used Padé approximation for the thermal ions, and finds that it can introduce a 10%  
 180 relative error near  $k_{\perp}\rho_c \sim 1.5$  comparing to the exact solution. Fig. 3 shows the comparison  
 181 between the four-point average method and the Padé approximation with the following form

182 
$$\Gamma_0(k_{\perp}\rho_c) = \frac{c_0}{1 + \frac{1}{c_0} \rho_c^2 k_{\perp}^2}. \quad (16)$$

187

183 We note that the Padé approximation with the Maxwellian distribution has a similar form with  
 184  $c_0 = 1$ . In the long wave length limit, these approximations are both very close to the exact  
 185 value, as is shown in Fig. 3, and it can be further verified by the numeric benchmarks shown  
 186 in the next section.

188

## IV. Numeric Verification

189 To verify our global algorithm for the slowing down background distribution, we shall solve  
 190 the gyrokinetic Poisson equation without electron response in a large-aspect-ratio toroidal  
 191 geometry with circular cross section as a sample problem. In the long wavelength limit, Eq.  
 192 (9) can be reduced to the following form using Taylor expansion of  $\Gamma_0$ :

$$193 \quad \frac{q_i}{m_i \Omega_i^2} \nabla_{\perp}^2 \delta\phi = -\frac{\delta n_{i,gc}}{n_i} \quad (17)$$

194 Toroidal effect can be ignored and  $\nabla_{\perp}^2 = \frac{1}{r} \frac{\partial}{\partial r} r \frac{\partial}{\partial r} + \frac{1}{r^2} \frac{\partial^2}{\partial \theta^2}$  in polar coordinates on poloidal  
 195 cross section. After these simplifications, Eq. (17) is just a normal Poisson equation and we  
 196 can choose  $\delta n_{i,gc}$  to be the eigenfunction of the Laplacian operator to ensure an analytic  
 197 solution. Let  $\delta n_{i,gc}/n_i = (J_m(k_0 r) - Y_m(k_0 r) J_m(k_0 a_1)/Y_m(k_0 a_1)) \cos m\theta$  in which  $k_0$   
 198 satisfies  $J_m(k_0 a_0) Y_m(k_0 a_1) - Y_m(k_0 a_0) J_m(k_0 a_1) = 0$ . Then the solution of this Poisson  
 199 equation is just  $\delta\phi = k_0^{-2} (J_m(k_0 r) - Y_m(k_0 r) J_m(k_0 a_1)/Y_m(k_0 a_1)) \cos m\theta$  with zero  
 200 boundary condition on  $r = a_0, a_1$ . With  $m = 6$  as an example, the comparison between  
 201 analytic solution and numeric solution along the line  $\theta = 0$  is shown on Fig. 3., where a  
 202 perfect match is found between them. More generally, the poloidal cross section contour for  
 203 the solution is shown on Fig. 4. The difference between the analytic 2D solution and numerical  
 204 one is negligibly small, as is shown by Fig. 4(c). Thus, in the long wavelength limit, our four-  
 205 point average method works perfectly for the slowing down equilibrium distribution.

206 In order to simulate short wavelength modes, we need to verify the validity of our  
 207 algorithm in the short wavelength limit. The verification process is getting subtle in the short  
 208 wave length limit, since there's no analytic solution for  $\delta\phi$  when expanding the  $\Gamma_0$  operator  
 209 in this limit. But we can still compare the numerical solutions to Eq. (12) by the developed  
 210 four-point average with the Padé approximation. We solve Eq. (9) with a natural unit  $T_i =$   
 211  $T_e = q_i = e = 1$ , and a short-wave-length density fluctuation  $\frac{\delta n}{n_i}$ , which is in the same form  
 212 as that in long-wave-length limit but with much larger  $m$  and  $k$ , i.e.,  $k_r \rho_c = 1 \sim 2, m =$   
 213  $62 \sim 125$ . As is shown in Fig. 5 and 6 for three cases with different  $k_r \rho_c$ , The solutions from  
 214 these two different numeric schemes show little difference, suggesting that they can both  
 215 handle the short wave length case within an acceptable error of 5%. The amplitude of the



216 solution using four-point average method is slightly larger, which can be ascribed to the fact  
217 that this operator of four-point average is larger than the Padé approximation in the  $k$  space,  
218 as is shown in Fig. 2.

219

220

## 221 V. Conclusion

222 A real space gyrokinetic Poisson solver for slowing-down equilibrium distribution has been  
223 developed based on the multi-point average method [2,16] and verified for its accuracy in  
224 the long and short wavelength limits. The discovery process for this method is shown in detail  
225 and it can be further modified to accommodate more equilibrium particle distributions. This  
226 method can be incorporated in the global gyrokinetic particle simulation to study the crucial  
227 alpha particle physics in the burning plasmas, i.e., to simulate the drift Alfvénic turbulence  
228 accurately in the presence of slowing-down alpha particle distribution.

229

230

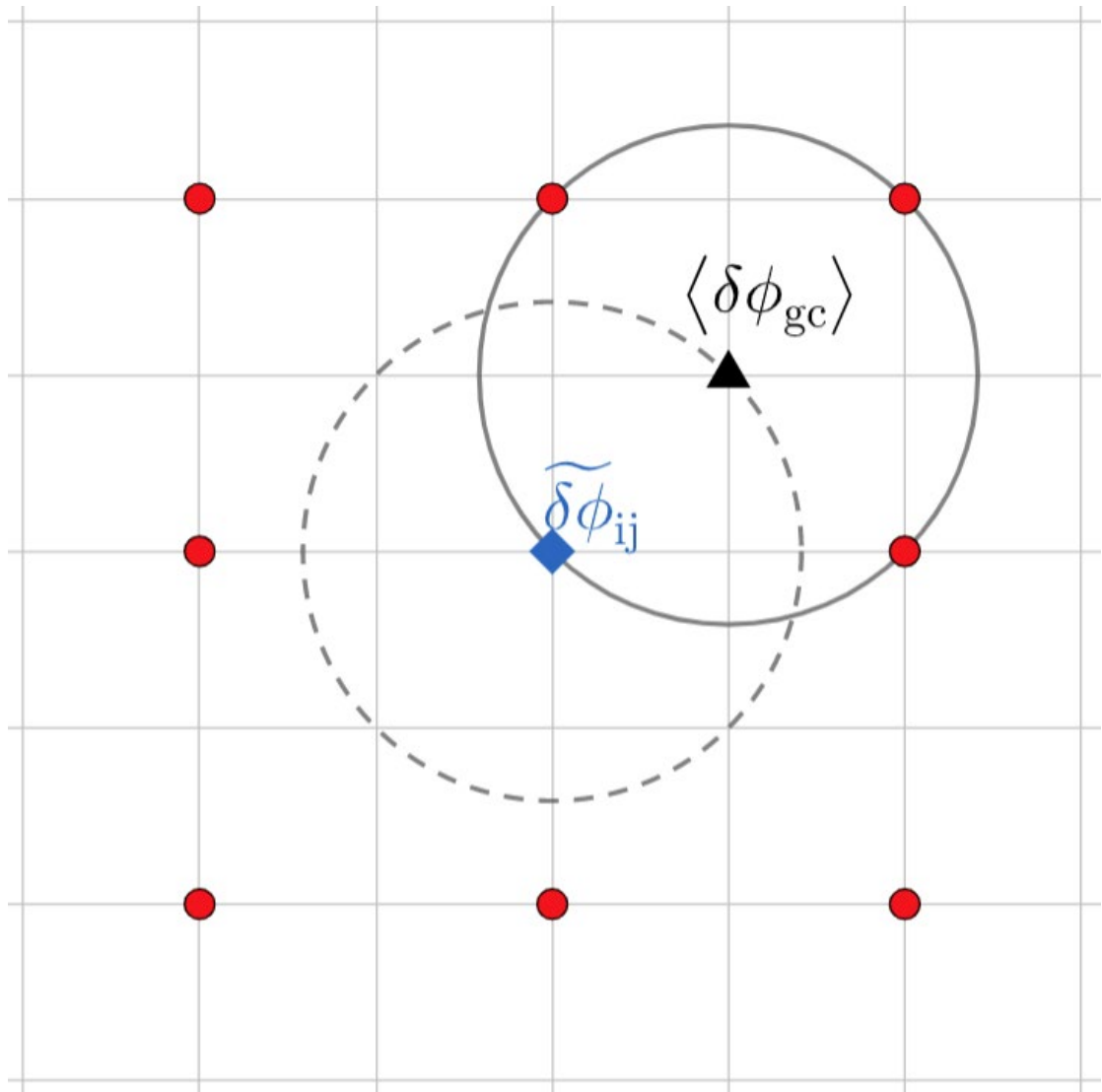
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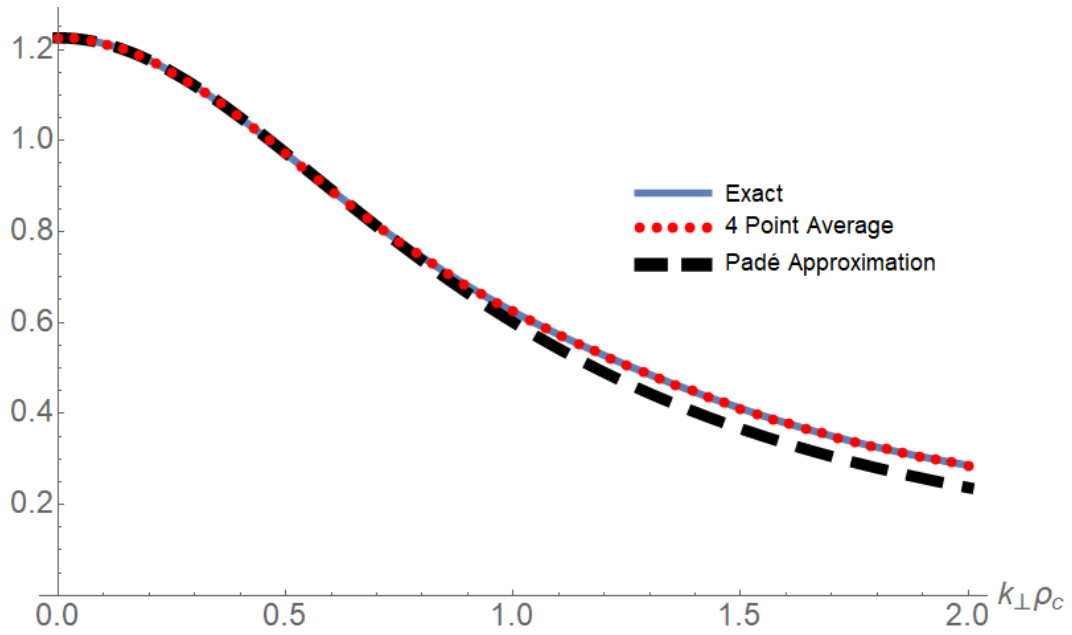
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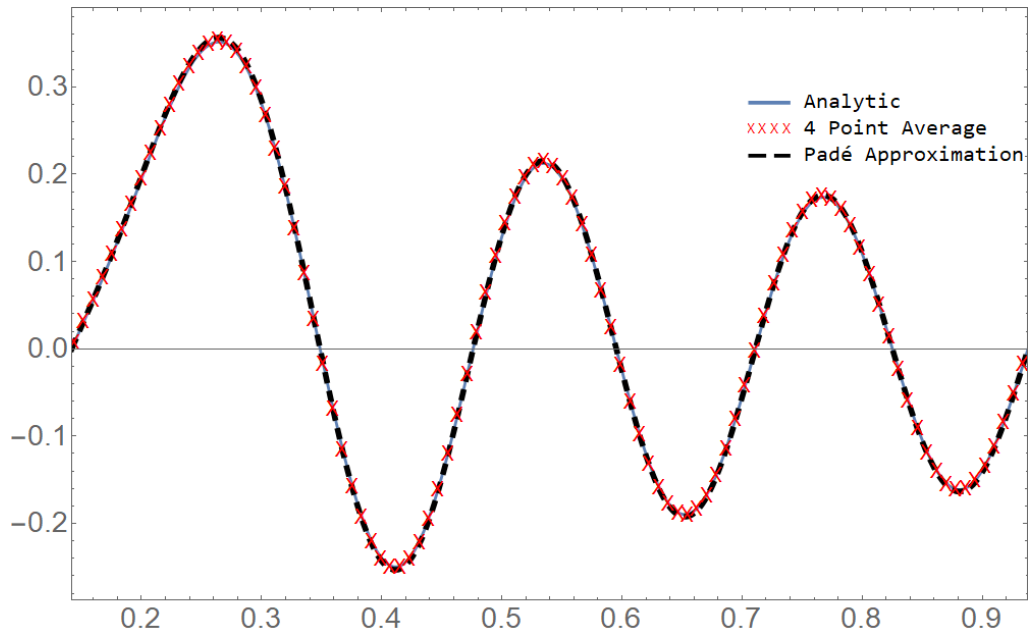
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Fig. 1. Scheme for calculating  $\tilde{\delta\phi}$  at grid point ij



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265 Fig. 2 Exact  $\Gamma_0$  function (blue solid line) and its numerical approximations vs  
 266 perpendicular wavelength  $k_{\perp}\rho_c$  : four-point average method with three velocity nodes in  
 267 the integration (red dotted line), and Padé approximation (black dashed line).



268

269

270 Fig. 3. Comparison of analytic expression and numeric solutions using 4-point average  
 271 approximation and Pade approximation for gyrokinetic Poisson equation in the long  
 272 wavelength limit with  $k_r\rho_i = 0.11$  and  $m = 6$ .

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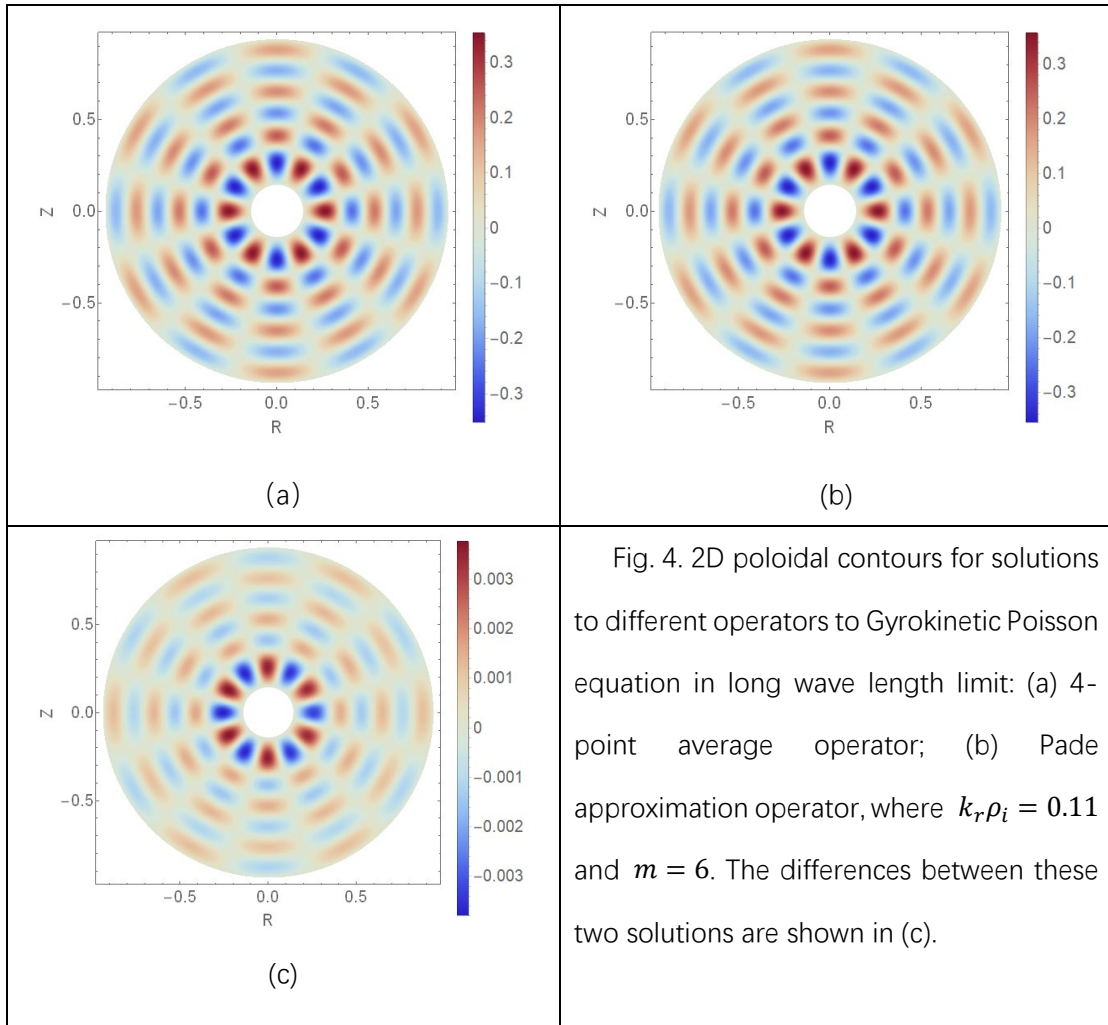
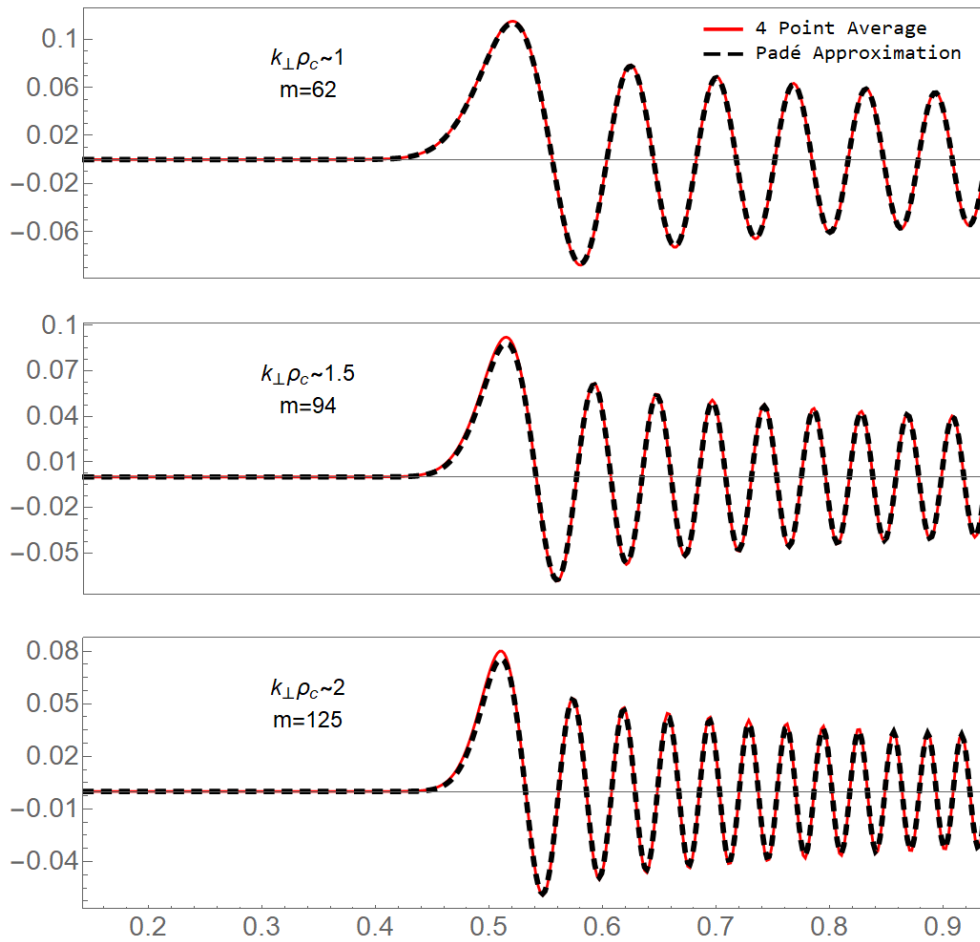


Fig. 4. 2D poloidal contours for solutions to different operators to Gyrokinetic Poisson equation in long wave length limit: (a) 4-point average operator; (b) Pade approximation operator, where  $k_r \rho_i = 0.11$  and  $m = 6$ . The differences between these two solutions are shown in (c).



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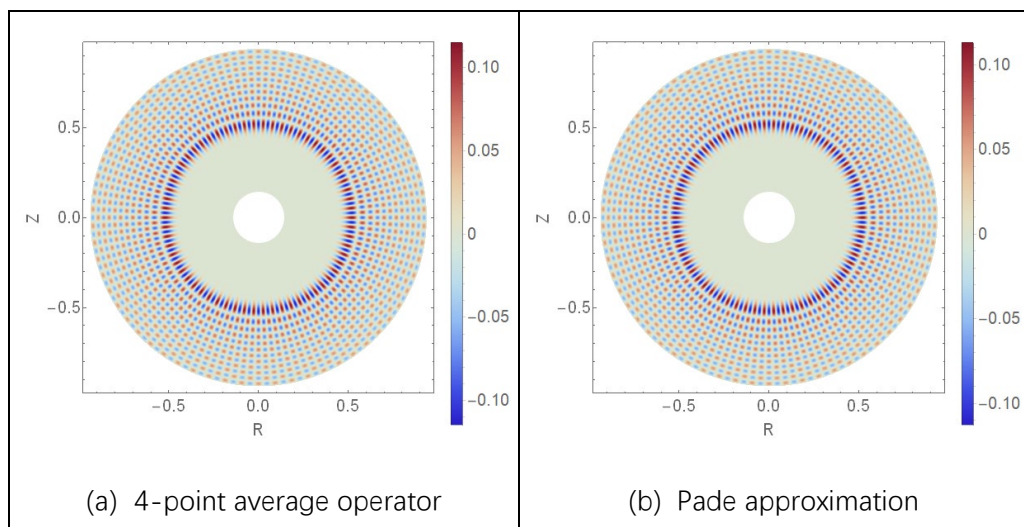
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Fig. 5. Comparison between 4-point average approximation (solid line) and Padé

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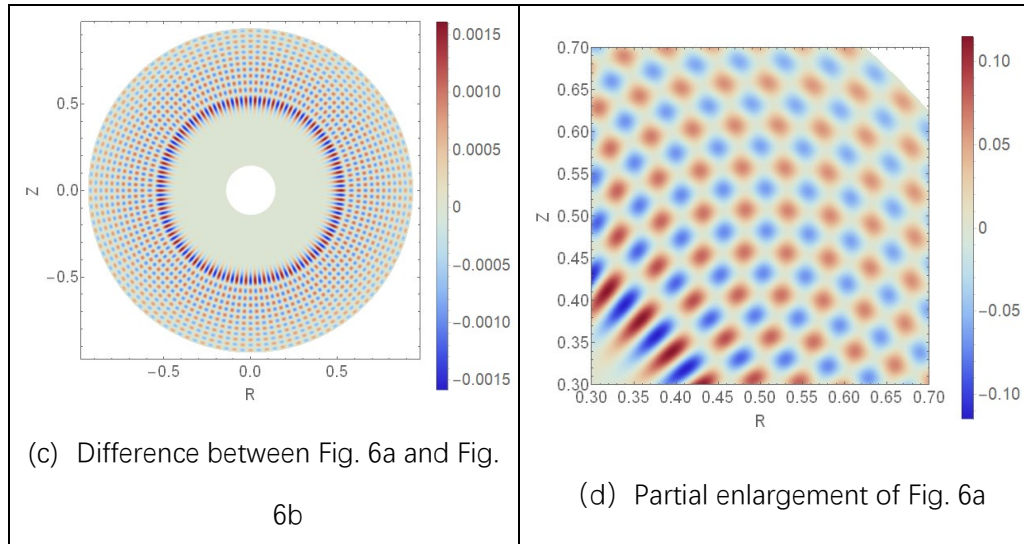
approximation (dashed line) for the solution to Poisson equation in the short wavelength limit.

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(a) 4-point average operator

(b) Padé approximation



279

280 Fig. 6. 2D contour of the solution to gyrokinetic Poisson equation in the short wave length  
 281 limit on the poloidal plane with  $k_r \rho_i = 1$  and  $m = 62$ . The numeric operator used in solving  
 282 the Poisson equation are (a) 4-point average operator, and (b) Pade approximation. The  
 283 difference between them is shown in (c), and the first quadrant of (a) is enlarged in (d) to  
 284 show its fine structure.

285